### Visualization and post-processing tools for $\operatorname{Siesta}$

Andrei Postnikov

Université Paul Verlaine, Metz

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# Outline

#### What to visualize?

2 XCrySDen by Tone Kokalj

#### 3 Sies2xsf suite

- Atomic structure
- Charge and spin densities
- Wave functions
- Fermi surfaces
- Molecular Dynamics or relaxation
- Phonons

- The structure (unit cell, positions of atoms)
- Charge (spin) density  $\rho(\mathbf{r})$ , or "local density of states": properties on the grid
- Kohn-Sham orbitals: properties expanded over the basis functions
- Fermi surfaces (or other isoenergy surfaces in **k**-space)
- Molecular dynamics or relaxation: how the atoms move (a movie)
- Phonon vibration modes (after a **Vibra/vibrator** run), shown by arrows or as a movie

#### Atomic structure

The error-free choice of structure (unit cell, positions of atoms) is the full responsibility is on the user; only minimal checks are done by SIESTA (e.g., "atoms too close").



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Since the input format is quite flexible (a big advantage!), it is difficult to organize a simple viewer of *input* structure data, without using the fdf routines. However, the .XV file (created after the completion of electronic structure loop) contains all necessary information, as it was *really* understood by SIESTA, in a fixed format: unit cell vectors and atom coordinates, all in Bohr. This file can be easily transformed to, e.g. .xyz format which is read by many vizualization programs (xmakemol, ...). In the following examples, we'll use XCrySDen.

# What to visualize..?

#### Charge/spin density; local density of states

These are scalar fields available, after a SIESTA run, on a 3-dim. mesh (the number of divisions along three lattice vectors is governed by the MeshCutoff parameter). A typical graphical representation of such scalar fields is by contour plots in 2-dim. cutting planes, and/or iso-



surfaces of a given level. Both representations are possible with XCrySDen.

#### Kohn-Sham wavefunctions

These are also scalar functions of spatial coordinates, but obtained in SIESTA as expansions over the basis functions. Their visual representation (as 2-dim. contours in a chosen plane, or 3-dim. isosurfaces) is handled by the **denchar** code. The **denchar** allows export of data in the Gaussian cube format, which can be read in by XCrySDen.

#### Fermi surfaces

can be calculated using the energy dispersion data  $E(\mathbf{k})$ , available from any band structure code. The difficulty of purely technical character is, how to construct energy isosurfaces and conveniently manipulate them (to choose viewpoint, select different sheets of the Fermi surface, etc.) This job is done within XCrySDen, provided the  $E(\mathbf{k})$  data are passed in a right format.





#### Molecular dynamics or relaxation

runs store the atomic positions in .MD and/or .ANI files (with and without unit cell information, correspondingly). Such sequences of atomic positions can be animated using various software packages, including XCrySDen.



#### Phonon modes

from Vibra/vibrator calculation after a

MD.TypeOfRun FC

SIESTA run can be represented by arrows (in a static figure), or as animations (a sequence of vibration snapshots).

# XCrySDen by Tone Kokalj, http://www.xcrysden.org



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# Format of XCrySDen input files (.xsf, .axsf, .bxsf)





# Format of XCrySDen input files (.xsf, .axsf, .bxsf)

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# Sies2xsf utilities

Where to find them:

- ··· siesta-2.0/Util/Contrib/APostnikov/ (old!)
- http://www.home.uni-osnabrueck.de/apostnik/Downloads, or mailto apostnik@uos.de or postnikov@univ-metz.fr

What do they make out of what:

- xv2xsf: .XV  $\rightarrow$  .XSF (view structure + cell)
- rho2xsf: .XV, .RHO or .LDOS → .XSF (atoms within a selected box + data grids)
- md2axsf: .XV, .ANI or .MD → .AXSF (animations of structure with fixed or variable cell)
- eig2bxsf: .XV, .KP and .EIG  $\rightarrow$  .BXSF (Fermi surfaces)
- vib2xsf: .XV and .vectors → .XSF and .AXSF for each selected phonon mode; static (with arrows to indicate dilacement patterns) and dynamic (animated phonon).

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# Visualization of atomic positions

xv2xsf asks for a .XV file and transforms it into .XSCF

I know my structure already, why visualize it?  $\quad \rightarrow \quad$  two reasons:

 xv2xsf uses .XV file, i.e. the structure information as is was understood by SIESTA, including possible input errors (messed up units etc.)
 xv2xsf allows to draw the simulation cell, which is useful in case of molecules or slabs: is there enough space around? Is it not too much? Note: XCrySDen draws all atoms *inside* the simulation box, irrespectively of their given positions. So you might want to see replicated cells ...



# Visualization of atomic positions

Example of for a crystal of Fe-binuclear units:



# Visualization of atomic positions



SnO<sub>2</sub> (rutile structure)

An unexpected bonus for SIESTA users: XCrySDen constructs the Brillouin zone from structure information in the .XCF file and allows to select k-path for plotting band structurse. Go to

 $\underline{T}ools \rightarrow k\text{-path Selection}$ 



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### Putting arrows on atoms

Suppose we have local (e.g., non-collinear) magnetic moments. Can we show them with XCrySDen?

 $\rightarrow$  Yes, we can use "Forces" entry in the XCrySDen input file format, columns #5-7. (But, there is no special tool for this. You should do it by hand, or write your own script). An example:

27	6.00000011	6.00000011	6.0000011	0.6220	0.6280	0.8430
27	7.24176814	7.94065215	6.00000011	1.1770	1.0700	0.8980
27	4.75823209	7.94065215	6.0000011	1.3540	0.7410	1.0070
27	7.24176814	4.05934808	6.0000011	0.4440	1.0860	1.4230
27	4.75823209	4.05934808	6.00000011	0.6370	0.7360	1.5460
27	6.00000011	7.24176814	7.94065215	1.1860	1.1880	0.7810
27	6.00000011	7.24176814	4.05934808	1.0440	0.4820	1.4290
27	6.00000011	4.75823209	7.94065215	0.7380	1.3350	1.0180
27	6.00000011	4.75823209	4.05934808	0.6020	0.6130	1.6410
27	7.94065215	6.00000011	7.24176814	0.8070	1.3540	0.9490
27	4.05934808	6.00000011	7.24176814	1.2290	1.0180	0.8930
27	7.94065215	6.00000011	4.75823209	0.5630	0.8130	1.5400
27	4.05934808	6.00000011	4.75823209	0.9880	0.4610	1.4830

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### Putting arrows on atoms

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#### Putting arrows where there are no atoms



XCrySDen allows to declare ghost atoms, labeled 'X'. They can be set very close to other atoms, so that XCrySDen won't complain – even at exactly the same place. The ghost atoms may have any radius, e.g. zero, and yet carry arrows.

One can make quite diferent use of such arrows. The only limitation: the properties of *all* arrows are fixed by

 $\underline{\mathsf{M}}\mathsf{odify} \to \mathsf{Force\ settings}$ 

and cannot be diversified.

# Visualization of Charge/spin densities, or of LDOS

- An input file for XCrySDen is created by **rho2xsf**.
- Accept the same approach as in Denchar: define the output box (by origin point and three spanning vectors, not necessarily orthogonal) and grid size along each grid direction. The values of a Siesta property defined on the internal Siesta grid are (linearly) interpolated onto the grid of the output box.
- The output box may be also 2-dimensional (No. of divisions =1 along one spanning vector).
- The output box may coincide with the Siesta box, or not. XCrySDen can apply translations to to generated grid.
- Ochoice isosurface parameters, cutting planes, isolines on the cutting planes, colors, lighting etc. from the means of XCrySDen.

# Visualization of Charge/spin densities, or of LDOS



# Visualization of Charge/spin densities, or of LDOS



Siesta simulation cell and grid

- $\bullet$  The fineness of the visualization is limited by the sparciest one of the two grids.
- The cutting planes in XcrySDen may only be those of the visualization grid.

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Loading the .XSF file  $File \rightarrow Open \ Structure \rightarrow Open \ XSF$ shows the atoms within the visualization box.



XCrySDen: AFM.XSI

This is a cut

(shown earlier)





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Similar, with isolines in a cutting plane. Cutting planes can only pass parallel to the edges of the visualization box. Therefore, the initial choice of the box (as rendered by **rho2xsf**) is important for a physically meaningful visualization.



Image: A matrix

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# Visualization of Wave Functions

In principle, **Denchar** does this job quite fine. However, we'd like to have an interface to XCrySDen. To this end:

- Use Denchar, define output box there, save result as Gaussian98 Cube file.
- **②** Read this Gausian98 Cube file into XCrySDen. Save in the XCrySDen format .xsf (  $\rightarrow$  A).
- A bug (or a feature?) in Denchar: it correctly translates the WF images (grid) over the output box, but not atoms.
- How to fix: run **rho2xsf**, define the same output box as in Denchar. Save the atom part in the XCrySDen format .xsf ( $\rightarrow$  B).
- Insert (by hand) the "correct" atom part from (B) into the place of "incomplete" atom part from (A).
- If needed, merge many grid blocks (which reside in different .xsf files, each exported from its own Gausian Cube) into a single .xsf file.

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The rest goes exactly as with the charge density. This example is for a fragment of a heterospin polymer (Cu ion in the complex with hexafluoroacetylacetonate, flanked by two free radicals).



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# Visualization of Fermi surfaces

 Calculate eigenvalues on a sufficiently fine, undicplaced k-mesh, i.e.

%block kgrid\_Monkhorst\_Pack

16 0 0 0.

0 16 0 0.

0 0 16 0. %endblock kgrid\_Monkhorst\_Pack

e get files .XV, .KP, .EIG, run eig2bxsf → creates .BXSF (or .BXSF\_1 and .BXSF\_2 for spin-polarized case).





# Visualization of Fermi surfaces

in XCrySDen, choose <u>File</u>  $\rightarrow$  Open Structure  $\rightarrow$  Open BXSF In the window that pops up, specify the Fermi energy



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# Visualization of Fermi surfaces

See which bands cross the Fermi energy, and select which of them you want to plot



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### Fermi surfaces of some elemental metals

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### Fermi surfaces of some elemental metals



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### Fermi surfaces of some elemental metals



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# Fermi surface of MgB<sub>2</sub>



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# Molecular Dynamics or Relaxation

Calculate MD history: WriteMDhistory T writes (updates existing) unformatted .MD file, either with or without variable cell; WriteMDXmol T writes (updates existing) formatted .ANI file (coordinates only, no variable cell information).

② Having .XV, and either .MD, or .ANI run md2axsf, answering questions about the (optional) choice of output box and the MD steps to visualize (first #; last #; keep only each #'s) → creates .AXSF.

Hopefully, variable or fixed cell will be recognized automatically. If cell information from .MD is not available, the .XV will be used (assuming fixed cell).

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- get .vectors (calculated by vibrator) and .XV (from Siesta)
- In vib2xsf, select # modes (first ... last) to vizualize. For each selected mode, a separate .XSF file and an .AXSF file are created. .XSF contains a static structures (as in .XV), with arrors added to each atom to indicate displacement pattern.

.AXSF contains the animation of a phonon, for a (user-chosen) amplitude and number of steps.

## An example of phonons in In- and N-doped GaAs



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# An example of phonons in In- and N-doped GaAs



In order to show

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# An example of phonons in In- and N-doped GaAs

One of N-related modes. The atoms shown are cut from much bigger supercell by selecting the "visualization box" in **vib2xsf**, and then some more atoms have been removed by hand.



# Concluding remarks and known limitations

- XCrySDen allows to manage (almost) all structure-related results from a SIESTA calcultion I can think about... do you have any other suggestions? It is a great software (both, I mean).
- The final results are bitmap (or, other format) damps of the contents of the simulation window (in the full-screen mode if needed), that would normally suffice for publication/presentation purposes. Atomic structure (atoms, bonds and cells, but not grid properties) can also be stored in [ <u>D</u>isplay → Lighting off ] mode as a postscript vector graphics.

# Concluding remarks and known limitations

- What is missing in XCrySDen is an option *to paint* an isosurface / Fermi surface *with a property* (e.g., Fermi velocity).
- Fermi-surface part of XCrySDen is not as flexible as real-space-grids part in what regards choosing view frame, colours, light sources etc.
- Making 2-dim. cuts of the Fermi surface is not implemented.
   However, this can be easily done by exporting the *E*(**k**) data ordered by the **eig2bxsf** script to other plotting routine.
- A possible extension: representing a vector field e.g., the magnetization m(r) from a non-collinear spin calculation. XCrySDen allows to draw a forest of arrors ("forces") stuck to ghost atoms of chemical label 'X', which atoms won't be visible if their radii set to almost zero, but arrors well visible is thick enough.

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